Acta Universitatis Carolinae. Mathematica et Physica

Pavel Mayer New UBV photometry of IU Aurigae

Acta Universitatis Carolinae. Mathematica et Physica, Vol. 29 (1988), No. 2, 83--97

Persistent URL: http://dml.cz/dmlcz/142594

Terms of use:

© Univerzita Karlova v Praze, 1988

Institute of Mathematics of the Academy of Sciences of the Czech Republic provides access to digitized documents strictly for personal use. Each copy of any part of this document must contain these *Terms of use*.



This paper has been digitized, optimized for electronic delivery and stamped with digital signature within the project *DML-CZ*: The Czech Digital Mathematics Library http://project.dml.cz

New UBV Photometry of IU Aurigae

PAVEL MAYER*)

Prague, Czechoslovakia.

Received 30 October 1987

The paper presents 356 UBV photoelectric measurements of the early-type eclipsing variable star IU Aurigae. The measurements were made in the years 1982—1985. The effect of the dynamical term on times of minima in a three-body system is discussed.

Jsou předložena nová fotoelektrická měření zákrytové proměnné raného typu IU Aurigae. 356 měření v systému UBV bylo získáno v letech 1982 až 1985. Je diskutován vliv dynamického členu na okamžiky minim v systému tří těles.

1. Introduction

Up to now, IU Aur (HD 35652; spectral type BO: p, magnitude in maximum V = 8.2, period $1.8115^{\rm d}$) is the only eclipsing variable the minima depths of which secularly change. Since the discovery of its variability (Mayer, 1964), the primary minimum depth in V colour has increased from $0.43^{\rm m}$ to $0.68^{\rm m}$. Most probably, the reason of this change is the rotation of the nodal line caused by a third body. More data about this system can be found in recent papers where light curves of this star are solved (Leung and Li, 1986; Drechsel and Mayer, 1987; Mayer and Drechsel, 1987). It is important to monitor the light curve changes, since then the parameters of the system can be determined more accurately.

In several papers photometry of IU Aur in different epochs has been obtained; these data are summarized in Tab. 1. In all cases, the comparison star was HD 35619, also of early spectral type, nearly identical colours, convenient brightness and only 3.7' distant (Table 2). Its UBV values by Landolt and Blondeau (1972) are based on 130 measurements from 6 nights, therefore they will be used throughput.

2. Measurements

In this paper, measurements obtained with the 65 cm telescope of the Charles University located in Ondřejov are presented. The measurements, numbering 356,

^{*)} Department of Astronomy and Astrophysics, Charles University, Švédská 8, 150 00 Praha 5, Czechoslovakia.

were made in 10 nights in the years 1982-85; they are given in Table 3 and plotted in Fig. 1.

In the years 1982 and 1983, the transformation equations to the standard UBV system were

$$\Delta V = \Delta v - 0.126 \Delta (B - V),$$

$$\Delta (B - V) = 1.190 \Delta (b - v),$$

$$\Delta (U - B) = 1.045 \Delta (u - b);$$

and in the years 1984 and 1985, they were

$$\Delta V = \Delta v - 0.078 \Delta (B - V),$$

$$\Delta (B - V) = 1.119 \Delta (b - v),$$

$$\Delta (U - B) = 1.020 \Delta (u - b).$$

Considering the colour differences following from data in Tab. 2 we see, that no transformation was necessary except in the deepest part of the primary minimum, where corrections not larger than 0.002^{m} were to apply.

The measurements were made mostly in order to precise the period of the variable, so the minima are covered well, but other parts of the light curve not. It appears that in maximum (however note that only the phase around 0.25 is covered) the variable is about 0.005^m brighter than in the year 1964. The change of brightness in minima is illustrated in Fig. 1.

The times of minima from the presented data were already calculated and have been published (Mayer, 1987). Normal points were also formed from the data and used for solving the light curve by the Wilson-Devinney code (Drechsel and Mayer, 1987; Mayer and Drechsel, 1987). The phases given in Tab. 3 have been calculated as already described in the latter paper: due to the light-time-effect amplitude being larger than 0.005^d , the phases, when closer than 0.1 (phase) to a minimum, were calculated using the measured minimum time, and for other measurements, the ephemeris

Pri.Min. = J.D.hel. $2438448 \cdot 4068 + 1 \cdot 811474^{d} E$ was used.

3. The dynamical term

The system of IU Aur consists of the eclipsing pair and a third body. The third body manifests by periodic behaviour of times of minima, known as the light time effect (LTE; see e.g. Irwin, 1959), and also by change of the orbital longitude of the close pair components. The largest amplitude of these changes has the term analogical to the annual equation in the motion of the Moon, with the period which equals to the orbital period of the third body. Exact formulae for the third body effects have been given by Söderhjelm (1982). The eclipsing pairs in IU Aur and similar systems

are close binaries with circular orbits. Therefore, we are interested in formulae for $e_1 = 0$; and, for simplicity, we will also consider $\gamma = 0$ and $C/G_2 = 1$ (in Söderhjelm's notation). Then, using Söderhjelm's formulae 5, 6 and 7, the formula for $\delta \lambda_1$ ($\lambda_1 = g_1 + h_1 + l_1$) can be written as

(1)
$$\delta \lambda_1 = 2k_1 \big[\big(\tfrac{2}{3} - Z \big) \big(\nu_2 - l_2 + e_2 \sin \nu_2 \big) + \big(Z - 2 \big) S \big] \,,$$
 where

$$k_1 = \frac{3}{8} \frac{m_3}{M} (1 - e_2^2)^{-3/2} \frac{P_1}{P_2},$$

$$Z = \cos i + \cos^2 i$$
,

$$S = \sin 2(v_2 + \omega_2) + e_2 \sin (v_2 + 2\omega_2) + \frac{1}{3} \sin (3v_2 + 2\omega_2).$$

Here j is the inclination between both orbits, P_1 the priod of the eclipsing pair, and the values with index 2 are the usual parameters of the third-body orbit. Approximately can be written $v_2 - l_2 + e_2 \sin v_2 = 3e_2 \sin v_2$. The effect on observed minima due to this dynamical term is

$$(O - C)_{\text{DYN}} = P_1 \cdot \delta \lambda_1 / 2\pi .$$

In case of coplanar orbits $\cos j = 1$, Z = 2 and the square parenthesis term equals $-4e_2 \sin v_2$. Note that when $e_2 = 0$ the dynamical term is zero only in the coplanar case; if $j \neq 0$, the amplitude can be even larger than if j = 0, since then the square parenthesis term equals $(Z - 2) \sin 2(v_2 + \omega_2)$.

The dynamical effect depends on P_1^2/P_2 , therefore is neglectably small for longer third-body periods. IU Aur presents a case where this effect can play a role.

The total effect on times of minima is

(3)
$$O - C = (O - C)_{LTE} + (O - C)_{DYN},$$

where

(4)
$$(O - C)_{LTE} = \frac{m_3}{M^{2/3}} \frac{P_2^{2/3} \sin i_2^0}{173(1 - e_2^2 \cos^2 \omega_2)^{1/2}} \left[\frac{\sin v_2}{1 + e_2 \cos (v_2 + \omega_2)} + e_2 \cos \omega_2 \right].$$

Periods are given in days; $M = m_1 + m_2 + m_3$.

There are 9 parameters in the formula (3): m_3 , M, P_1 , P_2 , j, i_2^0 , e_2 , ω_2 and T (implicitly in v_2). The periods and the total mass can be considered as known (or estimated); j and i_2^0 are joined by an equation with another angle (observable). From the shape of the O-C curve then remains to calculate 5 unknowns. This can be done e.g. by harmonic analysis or perhaps better by numerical optimalisation.

Another case where the dynamical term can play a role is FZ CMa (Moffat et al., 1983). The amplitude of the term can reach about 0.0005^d, and the mass function

can be only 70 to 80 % of its original value. Then the problem with too large mass of the third body, discussed by Moffat et al., is removed.

Given the present accuracy of the O-C curves in both cases (for IU Aur see Mayer, 1987, Fig. 2) it seems to have better to wait for some more minima times until the elaborated solution is attempted.

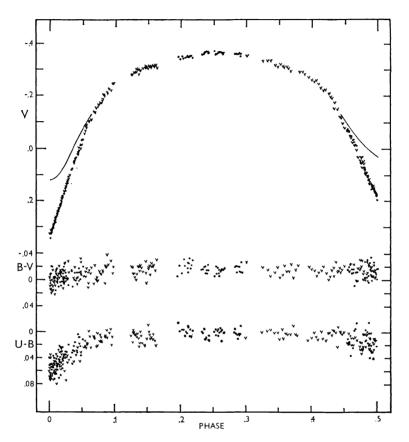


Fig. 1. V, B - V and U - B curves of IU Aur for the mean epoch 1984.4. To illustrate the change of minima depths, the V measurements from JD 2438667 (primary minimum) and JD 2439424 (secondary minimum) are schematically plotted.

References

- [1] DRECHSEL H., MAYER P.: Mitt. Astron. Ges. 68, 1987, 246.
- [2] EATON J. A.: Wisconsin Astrophys. No. 56, 1977.
- [3] HILTNER W. A.: Astrophys. J. Suppl. S. 2, 1956, 389.
- [4] IRWIN J. B.: Astronom. J. 64, 1959, 149.
- [5] JOHNSON H. L.: Ann. Astrophys. 18, 1955, 292.

- [6] LANDOLT A. U., BLONDEAU K. L.: Publ. Astron. Soc. Pacific 84, 1972, 394.
- [7] LEUNG K. C., LI Y. F.: Bull. Amer. Astron. Soc. 18, 1986, 977.
- [8] MAYER P.: Inf. Bull. Variable Stars No. 49, 1964.
- [9] MAYER P.: Bull. Astr. Inst. Czechosl. 22, 1971, 168.
- [10] MAYER P.: Bull. Astr. Inst. Czechosl. 27, 1976, 308.
- [11] MAYER P.: Bull. Astr. Inst. Czechosl. 34, 1983, 335.
- [12] MAYER P.: Bull. Astr. Inst. Czechosl. 38, 1987, 58.
- [13] MAYER P., DRECHSEL H.: Astron. Astrophys. 183, 1987, 61.
- [14] MOFFAT A. F. J., VOGT N., VAZ L. P. R., GRONBECH B.: Astron. Astrophys. 120, 1983, 278.
- [15] PAPOUŠEK J., VETEŠNÍK M.: Scripta Fac. Sci. Nat. Univ. Purkynianae Brun. Phys. 12, 1982, 435.
- [16] Rossati F.: Contr. Oss. Astr. Torino N. 42, 1967.
- [17] SÖDERHJELM S.: Astron. Astrophys. 107, 1982, 54.

Table 1. Published photoelectric measurements of IU Aur

Source	Mean epoch	Photometric system and number of measurements
Mayer (1971)	1964.5	U: 276; B: 250; V: 552
Rossati (1967)	1967.0	λ 515 nm: 592
Mayer (1971)	1968.7	U: 389; B: 424; V: 494
Mayer (1976)	1973.1	UBV: 420
Eaton (1977)	1974.9	λ 340 nm: 526; BV: 573
Papoušek and Vetešník (1982)	1976.6	U: 548; B: 756; V: 833
Mayer (1983)	1978.8	UBV: 112
This paper	1984.4	UBV: 356

Table 2. Data on the variable and comparison stars

Source/Phase	Johnson (1955) Hiltner (1956) Landolt and Blondeau (1972)	maximum secondary minimum primary minimum
U-B	-0.71 -0.69 -0.701	-0.699 -0.683 -0.644
B-V	$^{+0.24}_{+0.23}_{+0.241}$	+0.229 +0.230 +0.241
Δ	8.55 8.60 8.572	8.210 8.740 8.890
Declination 1950	+ 34°42.9′	+34°44.5′
Right ascension 1950	5 ^h 24 ^m 17 ^s	5 ^h 24 ^m 33 ^s
Catalog No.	HD 35619 BD+34°1046	HD 35652 BD+34°1051′
Star	Сошр.	IU Aur

Note: IU Aur UBV data are given for the epoch 1984.4, relatively to the HD 35619 values by Landolt and Blondeau (1972).

Table 3. Measurements of IU Aur (differences Var minus HD 35619)

J.D.hel.	ΔV	∆B	ΔU	Phase
2445240				
.5097	.152	.118	.155	.4904
.5180	.160	.151	.177	.4950
.5262	.196	.179	.200	.4995
.5336	.178	.163	.178	.5036
.5401	.150	.142	.142	.5072
.5481	.125	.114	.131	.5116
.5539	.103	.098	.101	.5148
.5200	.070	.055	.078	.5200
.5727	.040	.021	.040	.5252
.5789	.030	.012	.022	.5286
.5851	002	014	009	.5320
.5905	014	038	020	.5350
.5959	028	039	031	.5380
.6025	049	063	062	.5416
.6090	069	085	070	.5452
.6153	084	105	105	.5487
.6220	100	117	115	.5524
.6278	124	137	138	.5556
.6354	152	165	145	.5598
.6459	156	175	180	.5656
2445271				
.4283	119	138	141	.5562
.4334	146	160	170	.5590
.4389	— .157	186	187	.5620
.4437	— .176	182	185	.5646
.4495	190	206	202	.5679
.4543	200	206	207	.5704
.4597	20 9	208	208	.5734
.4661	220	231	220	.5770
.4730	220	241	232	.5808
.4794	234	250	252	.5844
.4867	248	260	250	.5884
.4928	259	266	260	.5918
.4990	257	272	273	.5951
.5061	261	278	271	.5991
.5152	274	295	283	.6041
.5432	286	309	305	.6196
.5503	295	312	310	.6234
.5558	293	298	309	.6264
.5633	297	322	320	.6306
.5707	307 -	321	318	.6347
.5764	310	324	323	.6378

Table 3 continued

J.D.hel.	ΔV	∆B	∆ <i>U</i>	Phase
.5813	310	- .332	323	.6406
.5862	316	325	330	.6433
.5919	320	325	315	.6470
.5999	309	328	328	.6508
.6085	318	330	329	.6556
.6196	336	340	344	.6618
.6276	336	—.351	354	.6662
.6417	336	356	— .357	.6740
.6859	350	376	—.368	.6983
445337				
.4016	.158	.170	.218	.9776
.4066	.182	.190	.222	.9804
.4093	.199	.183	.218	.9821
.4129	.213	.221	.250	.9841
.4155	.232	.225	.279	.9853
.4186	.229	.229	.282	.9870
.4210	.250	.245	.278	.9886
.4242	.250	.271	.317	.9903
.4270	.280	.279	.335	.9919
.4296	.282	.297	.335	.9932
.4322	.312	.296	.345	.9947
.4349	.301	.321	.353	.9962
.4376	.314	.326	.378	.9977
.4409	.310	.304	.369	.9995
.4439	.316	.308	.369	.0012
.4462	.320	.300	.367	.0025
.4487	.319	.301	.371	.0038
.4512	.299	.303	.339	.0052
.4550	.273	.273	.312	.0073
.4576	.264	.262	.307	.0088
.4600	.281	.266	.311	.0101
.4683	.223	.217	.261	.0147
.4708	.208	.213	.235	.0160
.4737	.213	.222	.246	.0176
.4794	.178	.157	.200	.0208
.4823	.150	.151	.188	.0224
445605				
.5033	.183	.184	.227	.9818
.5070	.212	.201	.248	.9838
.5114	.226	.228	.279	.9861
.5146	.243	.244	.291	.9880
.5192	.258	.266	.306	.9905

Table 3 continued

J.D.hel.	∆ V	∆B	ΔU	Phase
.5224	.271	.275	.321	.9923
.5270	.297	.301	.348	.9948
.5303	.317	.318	.367	.9967
.5352	.325	.333	.370	.9994
.5389	.321	.327	.373	.0014
.5435	.302	.311	.359	.0039
.5465	.296	.294	.346	.0056
.5513	.267	.265	.308	.0082
.5548	.243	.260	.300	.0100
.5589	.238	.242	.279	.0124
.5619	.226	.217	.257	.0141
.5667	.208	.200	.254	.0167
.5699	.190	.190	.222	.0185
.5753	.170	.154	.199	.0215
.5804	.145	.135	.175	.0243
.5855	.119	.118	.150	.0271
.5897	.103	.112	.133	.0294
2446004				
.4235	352	358	372	.1972
.4275	343	365	365	.1994
.4351	351	362	365	.2036
.4372	343	371	374	.2048
.4421	347	368	362	.2075
.4448	351	375	375	.2090
.4475	354	371	369	.2104
.4 498	352	363	372	.2117
.4519	348	371	378	.2129
.4542	349	378	371	.2141
.4564	355	372	375	.2154
.4587	356	376	369	.2166
.4611	347	374	376	.2180
.4882	359	374	374	.2329
.4918	363	378	376	.2349
.4938	362	380	384	.2360
.4957	365	384	372	.2371
.4988	373	382	371	.2388
.5007	369	383	374	.2398
.5028	370	379	367	.2410
.5047	362	383	370	.2420
.5069	369	383	377	.2432
.5240	372	377	377	.2527
.5260	358	—.377	372	.2538
.5280	363	371	374	.2549

Table 3 continued

J.D.hel.	∆ V	∆B	ΔU	Phase
.5307	361	379	380	.2564
.5327	370	384	—.377	.2575
.5348	367	376	379	.2586
.5368	370	375	373	.2597
.5387	370	381	386	.2608
.5408	370	384	384	.2619
.5428	366	381	367	.2631
.5450	364	377	380	.2643
.5475	360	374	373	.2656
.5498	367	383	382	.2669
.5793	365	383	382	.2832
.5813	363	383	387	.2843
.5834	361	372	369	.2855
.5863	357	376	371	.2870
.5882	361	379	368	.2881
.5904	365	377	373	.2893
.5962	363	374	369	.2905
.5947	353	376	372	.2917
.5969	362	372	380	.2929
2446005				
.5799	312	341	320	.8356
.5819	322	335	323	.8367
.5870	315	337	335	.8395
.5888	321	339	329	.8405
.5913	316	335	328	.8419
.5933	315	339	331	.8430
.5963	313	332	322	.8446
.5983	317	325	322	.8457
.6005	313	—.333	335	.8469
.6027	307	324	337	.8482
.6082	309	305	297	.8512
.6104	316	326	306	.8524
.6126	310	337	318	.8536
.6183	308	321	305	.8568
.6210	298	316	311	.8583
.6231	301	320	322	.8594
.6251	295	314	298	.8605
.6273	302	316	295	.8617
.6293	28 4	302	294	.8628
.6313	302	321	320	.8639
.6335	293	301	295	.8652
.6363	302	303	300	.8667
.6384	291	296	293	.8679

Table 3 continued

J.D.hel.	∆ V	∆B	∆ <i>U</i>	Phase
.6408	292	2 96	2 97	.8692
.6429	290	299	304	.8703
.6454	280	— .297	297	.8717
.6473	285	294	280	.8728
.6493	283	302	298	.8739
2446018				
.3836	244	263	255	.9018
.3869	242	251	—.231	.9037
.3905	221	248	253	.9057
.3956	228	250	249	.9085
.3985	209	244	235	.9101
.4006	212	230	232	.9112
.4036	214	231	225	.9129
.4058	203	199	191	.9141
.4103	203	214	198	.9166
.4123	188	202	188	.9177
.4150	183	194	185	.9192
.4171	180	188	174	.9203
.4194	175	185	181	.9216
.4216	165	178	164	.9228
.4257	165	164	162	.9251
.4278	167	172	152	.9262
.4297	161	161	138	.9273
.4319	156	158	148	.9285
.4338	138	160	152	.9296
.4362	—.141	136	130	.9309
.4448	113	131	107	.9356
.4467	111	121	114	.9367
.4491	099	113	—.107	.9380
.4515	106	097	098	.9393
.4535	102	099	079	.9404
.4555	090	101	084	.9415
.4575	078	095	070	.9426
.4596	083	089	067	.9438
.4615	069	082	069	.9449
.4641	—.057	064	040	.9463
.4707	034	059	028	.94 99
.4727	—.015	044	001	.9510
.4746	016	027	012	.9521
.4772	011	022	006	.9535
.4791	013	012	008	.9546
.4812	005	009	.014	.9557
.4832	009	002	.027	.9568

Table 3 continued

J.D.hel.	<i>∆V</i>	∆B	∆ <i>U</i>	Phase
.4850	.009	.006	.039	.9578
.4869	.014	.014	.046	.9589
.4899	.035	.017	.061	.9605
.4920	.039	.036	.066	.9617
.4940	.047	.044	.088	.9628
.4960	.052	.040	.087	.9639
.5061	.097	.086	.127	.9695
.5080	.116	.090	.135	.970
.5101	.114	.112	.147	.9717
.5126	.125	.123	.160	.973
.5147	.122	.111	.183	.9742
.5168	.141	.118	.117	.9754
.5259	.182	.171	.217	.9804
.5283	.192	.185	.229	.9817
.5308	.200	.191	.243	.9831
.5336	.224	.204	.272	.984
.5369	.229	.225	.303	.9859
.5505	.280	.277	.348	.9940
.5528	.302	.318	.366	.995
.5559	.313	.327	.366	.9970
.5583	.320	.316	.375	.9983
.5607	.318	.314	.381	.999
.5634	.342	.334	.406	.001
.5732	.296	.290	.333	.006
.5765	.283	.271	.344	.008
.5787	.263	.247	.297	.009
.5812	.255	.246	.308	.010
.5834	.239	.231	.291	.012
.5859	.236	.247	.275	.013
.5887	.216	.212	.250	.015
.5928	.200	.191	.230	.017
.5953	.189	.194	.225	.018
.5975	.173	.173	.230	.020
.6001	.165	.171	.208	.021
.6026	.152	.149	.197	.022
.6050	.149	.125	.165	.024
.6225	.079	.066	.106	.033
.6245	.071	.059	.092	.034
.6271	.044	.054	.084	.036
.6294	.047	.047	.070	.037
.6315	.033	.024	.055	.038
.6337	.032	.034	.051	.039
.6359	.030	.011	.044	.041
.6388	.019	.009	.042	.042

Table 3 continued

J.D.hel.	∆ V	∆B	<i>∆U</i>	Phase
.6410	.012	.002	.007	.0439
.6430	.006	010	.000	.0450
.6452	017	017	007	.0463
.6473	018	026	002	.0474
.6493	034	041	005	.0485
.6548	037	054	041	.0516
2446335				
.5225	.135	.130	.171	.9747
.5259	.154	.155	.190	.9770
.5313	.180	.167	.218	.9797
.5350	.199	.193	.230	.9818
.5392	.216	.212	.256	.9840
.5435	.234	.233	.276	.9864
.5467	.245	.240	.298	.9882
.5509	.264	.260	.316	.9904
.5545	.276	.272	.330	.992
.5582	.283	.289	.352	.994
.5621	.304	.309	.370	.996
.5654	.316	.320	.368	.998
.5687	.316	.316	.368	.0004
.5722	.310	.316	.364	.002
.5760	.290	.300	.358	.004
.5815	.274	.281	.336	.0074
.5857	.271	.265	.318	.0098
.5894	.250	.235	.300	.0118
.5932	.230	.219	.272	.0138
.5974	.208	.206	.250	.0162
.6015	.203	.196	.242	.0184
.6059	.187	.170	.195	.0209
.6127	.145	.129	.160	.0248
2446363				
.5623	077	090	081	.4553
.5641	069	086	066	.4563
.5665	063	082	062	.4577
.5681	065	085	069	.4585
.5700	049	062	058	.4596
.5719	052	066	030	.4606
.5758	048	0 57	036	.4628
.5773	035	049	023	.4636
.5811	027	018	022	.4657
.5826	0 19	017	023	.4665
.5843	016	013	028	.4675

Table 3 continued

J.D.hel.	∆ V	∆B	∆ <i>U</i>	Phase
.5860	.001	008	.001	.4684
.5879	013	011	.003	.4695
.6093	.080	.056	.068	.4813
.6107	.081	.068	.063	.4821
.6127	.085	.074	.085	.4832
.6143	.088	.064	.080	.4840
.6159	.099	.072	.070	.4849
.6176	.096	.085	.093	.4859
.6191	.099	.103	.087	.4867
.6206	.123	.118	.119	.4875
.6222	.120	.097	.105	.4884
.6236	.130	.116	.123	.4892
.6256	.124	.117	.135	.4903
.6272	.134	.135	.143	.4912
.6288	.133	.121	.142	.4921
.6303	.143	.130	.157	.4929
.6318	.149	.130	.160	.4937
.6334	.150	.139	.154	.4946
.6350	.161	.135	.174	.4955
.6382	.157	.148	.167	.4972
.6398	.163	.155	.167	.4981
.6417	.167	.166	.187	.4992
.6433	.169	.163	.171	.5001
.6450	.163	.160	.172	.5010
.6466	.167	.152	.169	.5019
.6484	.163	.148	.156	.5029
.6501	.167	.146	.173	.5038
.6518	.151	.146	.176	.5047
.6536	.139	.131	.150	.5057
.6552	.139	.133	.153	.5066
.6570	.136	.119	.147	.5076
.6584	.133	.115	.138	.50 84
.6600	.127	.115	.131	.5093
.6620	.130	.112	.138	.5104
.6635	.130	.108	.129	.5112
.6650	.112	.099	.110	.5120
.6666	.105	.095	.096	.5129
.6682	.104	.105	.104	.5138
.6731	.080	.076	.0 96	.5165
.6748	.083	.082	.083	.5174
.6764	.075	.075	.076	.5183
.6780	.068	.042	.063	.5192
.6797	.053	.044	.070	.5201
.6 816	.043	.041	.056	.5212

Table 3 continued

J.D.hel.	∆ <i>V</i>	∆B	∆ U	Phase
.6832	.066	.037	.068	.5221
.6854	.048	.029	.049	.5233
.6870	.050	.022	.046	.5242
.6893	.025	.009	.023	.5254
2446364				
.5038	.130	.132	.169	.9744
.5077	.154	.156	.193	.9766
.5110	.170	.166	.204	.9785
.5144	.176	.170	.224	.9802
.5177	.182	.187	.232	.9821
.5211	.198	.200	.248	.9840
.5247	.220	.222	.267	.9860
.5282	.228	.232	.291	.9880
.5316	.248	.249	.304	.9898
.5350	.264	.251	.314	.9917
.5390	.284	.278	.328	.9937
.5450	.300	.305	.370	.9966
.5487	.308	.308	.366	.9993
.5536	.304	.302	.368	.0020
.5568	.298	.306	.360	.0037
.5609	.275	.279	.342	.0060
.5643	.267	.265	.308	.0082
.5676	.256	.258	.308	.0096
.5710	.241	.238	.294	.0116
.5746	.230	.234	.276	.0136
.5794	.202	.202	.249	.0161
.5841	.178	.166	.233	.0187
.5877	.164	.156	.210	.0208
.5911	.146	.151	.192	.0226
.5943	.128	.125	.162	.0244